HONORS TRACK RADIO ASTRONOMY

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Content of today

- Who are we?
- Radio Astronomy
- Structure of the module
 - Seminars
 - What are you supposed to do ...
- Questions

Mark Bentum

- TU/e Electrical Engineering : Professor Radio Science & Dean
- ASTRON was Head of the Astronomy & Operations department

TU/e AST(RON Netherlands Institute for Radio Astrono







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Jasmina Lazendic-Galloway

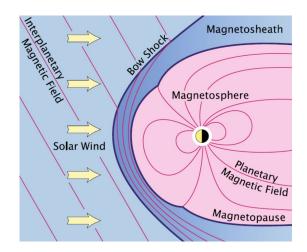


Simon Petyniak (BSc Applied Maths)

HA RA project: Star-planet interactions

Magnetospheres and Habitability

- Separates the regions dominated by IMF and planetary magnetic field
- Helps maintain atmosphere and liquid water on the surface (McIntyre et al, 2019)
- Protects the planet from radiation



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Goals

- Your personal development
- Learn about astronomy / cosmology / life origins
- Learn to study complex matter
- Learn to work & communicate in teams
- Work on a large challenge-based project
- You are in the driving seat and we are here to help:
 - Both applied and research-based projects are possible
 - Non-physics majors are also very welcome (we have computer science and architecture students! Everyone's expertise can be applied to astronomy or space projects!)

Question you might have ...

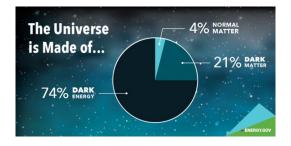
- How did our universe come into existence?'
- Is life possible in other places in our galaxy?
- What is intelligent behavior?
- What is a black hole?
- What is the faith of our planet, our solar system, our galaxy, our universe ??'

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The 5 biggest questions about the Universe

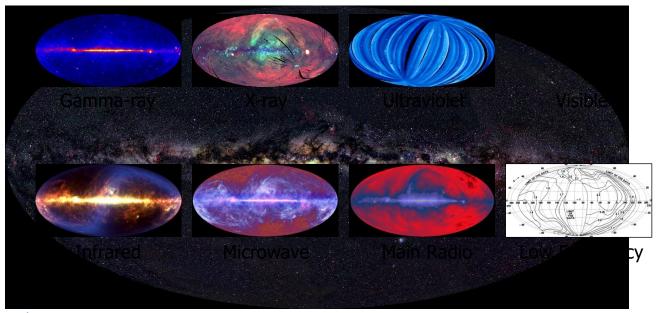
- What is dark matter?
- What is dark energy?
- What came before the big bang?
- What's inside a black hole?
- Are we alone?





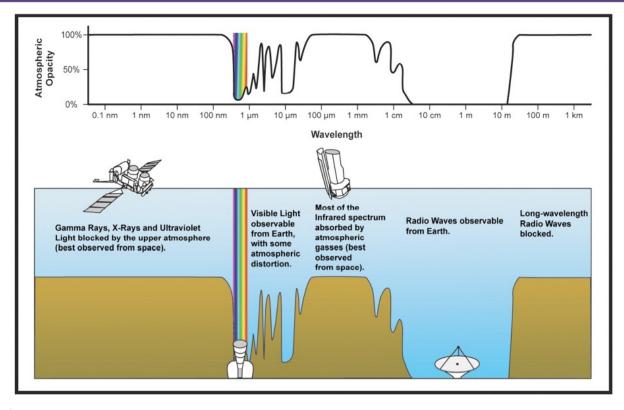
"Images" from space

- Celestial sky: has been mapped for nearly every type of electromagnetic radiation
- major exception: ultra-low frequency radiowaves (<30 MHz)



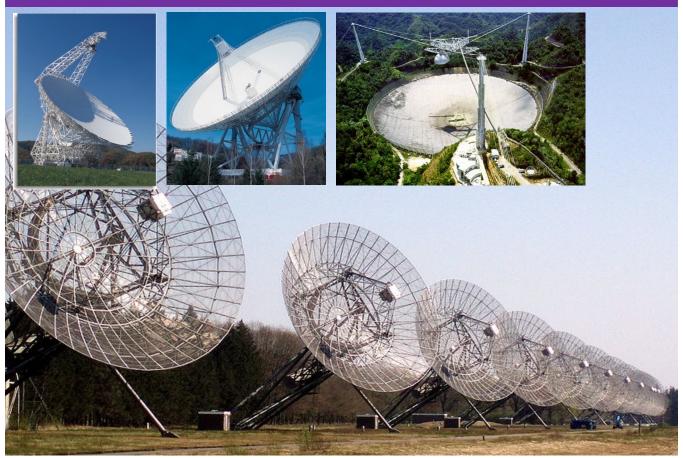
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How to capture the signals



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Radio



Example: Supernova Remnants across λ

- Radio: Accelerated particles: synchrotron radiation from e⁻ in magnetic field
- > Optical: Continuum and emission lines from ejecta, progenitor wind, blast wave
- > Infrared: Continuum and emission lines from cooling ejecta, shock-heated dust
- X-ray: Continuum and emission lines from shock wave, ejecta, accelerated particles, compact central object and pulsar wind nebulae
 Increasing energy

0.0001 nm 0.01	nm 1		ing wavelength 000 nm 0.01 cm	1 cm	1 m	100 n
Gamma rays	X-rays	nys Ultra- violet Infrared		Radio waves		
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Interesting topics (not limited)

- Navigation using pulsars
- Making a radio telescope
- Al system for searching for features in radio astronomy data
- Auxiliary science with astronomical data (example: lightning)
- > SETI
- Cosmic-ray acceleration
- Star-planet interactions

Structure of this module

(to be finalized ... also depending on you)

- A. Seminars to introduce you to the topics
- B. Reading books / studying articles etc..
- c. Working on a project (CBL)
 - A. Several groups ..
 - B. For two years (two one-year long projects are also possible)
- D. Of course: visit to ASTRON Dwingeloo/Westerbork/ LOFAR/..
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Application

- Due 12 May
- If you have more questions:
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- Jasmina: j.lazendic.galloway@tue.nl
- HA RA student Simon: s.w.petyniak@student.tue.nl